

# General Relativity

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## 1 General Relativity

### 1.1 Why Force is Like Curvature

- Alice and Bob on the earth's equator, at a large yelling distance away from each other
- Earth is frictionless
- Pushed straight north
- Eventually will impact each other as the great circles they move along collide with each other
- Possibilities

**Force** Bob thinks his path is being curved by some outside force

$$F_0 = -m \left( \frac{v}{B} \right) l(t) \quad (1)$$

$$B = 6400KM \quad (2)$$

**Curvature** Alice thinks they are travelling along a curved surface

$$\text{radius of curvature} = R_c = B \quad (3)$$

$$\theta = \frac{l_0}{r_0} = \frac{l_0}{R_E} = \frac{l(t)}{r(t)} \quad (4)$$

$$F = ma = m \frac{d^2}{dt^2} l(t) \quad (5)$$

$$r(t) = R_E \cos(\phi(t)) \quad (6)$$

$$l(t) = \frac{l_0}{R_E} r(t) \quad (7)$$

$$= \frac{l_0}{R_E} R_E \cos(\phi(t)) \quad \text{substitute in } r(t) \quad (8)$$

$$= l_0 \cos \left( \frac{vt}{R_E} \right) \quad \text{Cancel and substitute in } \phi(t) \quad (9)$$

$$\phi(t) = S/R_E = vt/R_E \quad (10)$$

$$F = -m \left( \frac{v}{R_E} \right)^2 l(t) \quad (11)$$

## 1.2 Equivalence Principle

- It is always possible, for some object and force, to replace a force by some equivalent curvature
- For gravity, this works particularly well

$$m_i = \text{inertial mass} \tag{12}$$

$$m_g = \text{gravitational mass} \tag{13}$$

$$F = m_i a \tag{14}$$

$$F_N = - \left( \frac{GM}{r^2} \right) m_g \tag{15}$$

$$m_i a = \frac{GM}{r^2} m_g \tag{16}$$

$$a = \left( \frac{GM}{r^2} \right) \frac{m_g}{m_i} \tag{17}$$

$$m_g = m_i \tag{18}$$

$$\tag{19}$$

All bodies fall with equal acceleration

- GR = SR + Newton's Law + Straight Lines

## 2 Topology and General Relativity

### 2.1 Topology

- Intuition: Vectors — more than just numbers, has other invariants even if you try to change the coordinates by rotating (eg, length)
- Imagine a surface like a sphere

$$R^2 = x^2 + y^2 + z^2 \tag{20}$$

$$x' = zx \tag{21}$$

$$y' = y/2 \tag{22}$$

$$z' = 3z \tag{23}$$

$$R^2 = \left( \frac{x'}{2} \right)^2 + (2y')^2 + \left( \frac{z'}{3} \right)^2 \tag{24}$$

Distorting the coordinates give you an ellipsoid

- Invariants

**Handles** donut, coffee cup, torus all have one (basically a big hole that you could grab on to with

**Holes** If you remove one point from a sphere, you get something equivalent to a plane

**Twists** Möbius strip has one

### 2.2 Requirements for GR

- All observers are equal (as long as gravity is the only force in effect)
- Alternatively, all coordinate systems are equivalent
- Hence, need things that are invariant with respect to coordinate system
  - Topology

- Events
- What is the difference between a donut and a coffee cup? Geometry
  - Straight lines (or just distances) (metric =  $g$ )
  - Derivatives (connection =  $\Gamma$ )
- In GR,  $g \leftrightarrow \Gamma$ , but not in quantum gravity
- What you need for GR:
  - Topology
  - Matter
  - Coordinate system

Plug into Einstein equations, get a metric  $g$

### 3 Black Holes

#### 3.1 Schwarzschild Black Hole

- Stuff for GR

**Topology**  $\mathbb{R}^4 - \text{line}$  (three space minus the point for the singularity, crossed with a line for time:  
 $(\mathbb{R}^3 - \{\text{point}\}) \oplus \mathbb{R}^1$ )

**Matter** Spherically symmetric

**Coordinate system** Far away from matter (fiducial observer = “Fido”)

- Get metric

$$ds^2 = - \left(1 - \frac{2M}{r}\right) dt^2 + \left(\frac{1}{1 - \frac{2M}{r}}\right) dr^2 + \dots \quad (25)$$

- Meaning

- $ds$  = space-time distance
- $dr$  = physical distance
- $dt$  = time distance

- Problems

- $r = 0$ : always an issue
- $r = 2M$ : only because of coordinates

#### 3.2 Sonic Black Hole

- Take a huge pond
- Drill a hole in the bottom, with a spike in the bottom
- Water flows out of the hole faster than the speed of sound in the water
- Take tadpoles in the pool

**Freefo** Free fall observer

**Fido** Fiduciary observer

- Able to “see” solely using sonar
- Equipped with a regular noisemaker (eg, loud heartbeat)
- As Freefo falls in, Fido hears Freefo’s heart slowing down
- Past the horizon where the water starts moving faster than the speed of light, Freefo’s heart beat (which travels as waves in the water) will never get out
- From Freefo’s point of view, the horizon isn’t anything too special

## 4 Semi-classical GR

- Hawking radiation
  - $\Delta x \Delta p = \hbar$
  - Near the black hole, uncertainty about where exactly things are
  - Because particles spontaneously appear and disappear, some energy will escape from the black hole
  - Eventually the black hole will completely evaporate (maybe, if insufficient matter gets sucked in)
  - Black holes have a “temperature”:  $T_{BH} = \frac{\hbar c^3}{8\pi G M k_B}$  is proportional to  $\frac{1}{M}$
  - Information entropy
    - \* Disorder
    - \* Classical: Black hole is perfectly ordered, entropy = 0
    - \* Quantum: Black hole has screwy boundary:

$$\text{entropy} = S = \text{Information (bits)} = \frac{1}{4} \left( \frac{A}{l_{pl}^2} \right) \quad (26)$$

$$l_{pl} = \text{scale where spacetime loses its meaning} \quad (27)$$

$$\Delta x \Delta p = \hbar \quad (28)$$

$$\Delta p = p \quad (29)$$

$$\Delta x = l \quad (30)$$

$$\therefore pl = \hbar \quad (31)$$

$$\therefore p = \hbar/l \quad (32)$$

$$p = \frac{E}{c} = \frac{m_g c^2}{c} = m_g c \quad (33)$$

$$m_g = p/c = \frac{\hbar}{ec} \quad (34)$$

$$l = \frac{R_{swh}}{2} = \frac{Gm_g}{c^2} = \frac{G\hbar}{lc^3} \quad (35)$$

$$\text{Multiplying by } l \text{ and taking the root, } l = \sqrt{\frac{G\hbar}{c^3}} = l_{pl} \quad (36)$$

$$(37)$$

So, GR and quantum act strangely together, because we can randomly get black holes

- Now we get things quantized at the size of  $l_{pl}$ , so we have just one bit per small region, instead of one bit per point

## 5 Quantum Gravity and Holography

### 5.1 Holography

- All these black holes reduce our freedom
- Final entropy:

$$S_f = \frac{1}{4} \frac{A}{l_{pl}^2} \quad (38)$$

- By the second law of thermodynamics,

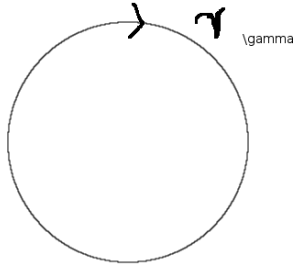
$$S_i \leq S_f = \frac{1}{4} \frac{A}{l_{pl}^2} \quad (39)$$

- So, the information is proportional to the *surface area*, so is the world 2D?

### 5.2 Quantum Gravity

- Look at (2D space + 1D time = 3D spacetime)
- This has been solved
- Observables = integrals over path in spacetime

$H[\gamma] = \text{knots}$



- Quantum gravity in 2 space dimensions is holographic

### 5.3 String Theory

- Lives on a surface in 10 dimensions (4 space dimensions, 1 time dimension, plus a 5 sphere at every point in space)
- Drop a dimension, because we look at the boundary (3+1 dim)
- Correspondence between String theory and Conformal Field Theory